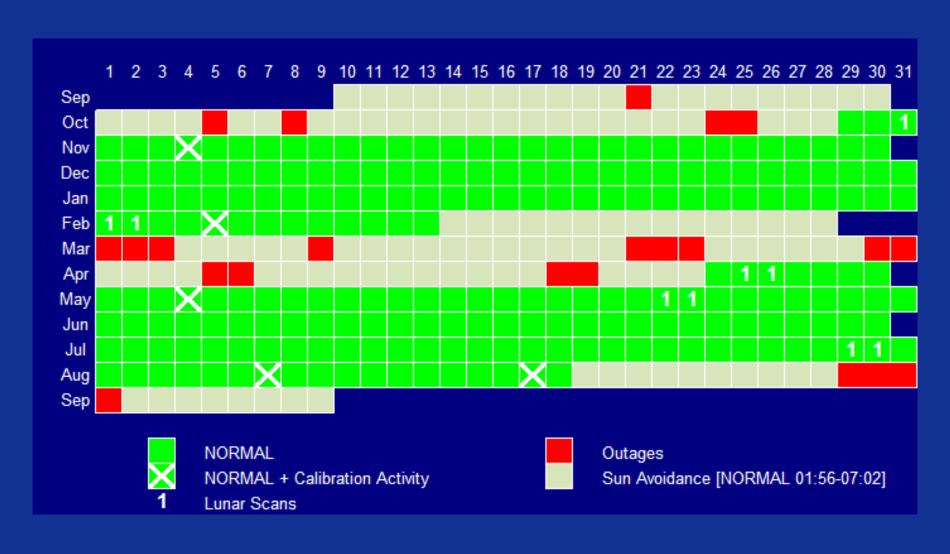
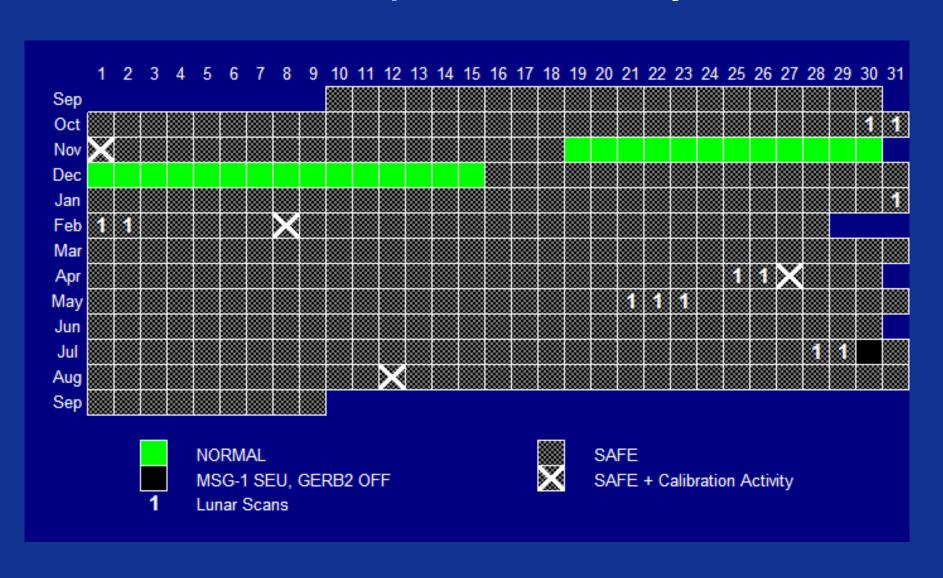
GERB Operations: Status Report J. Rufus

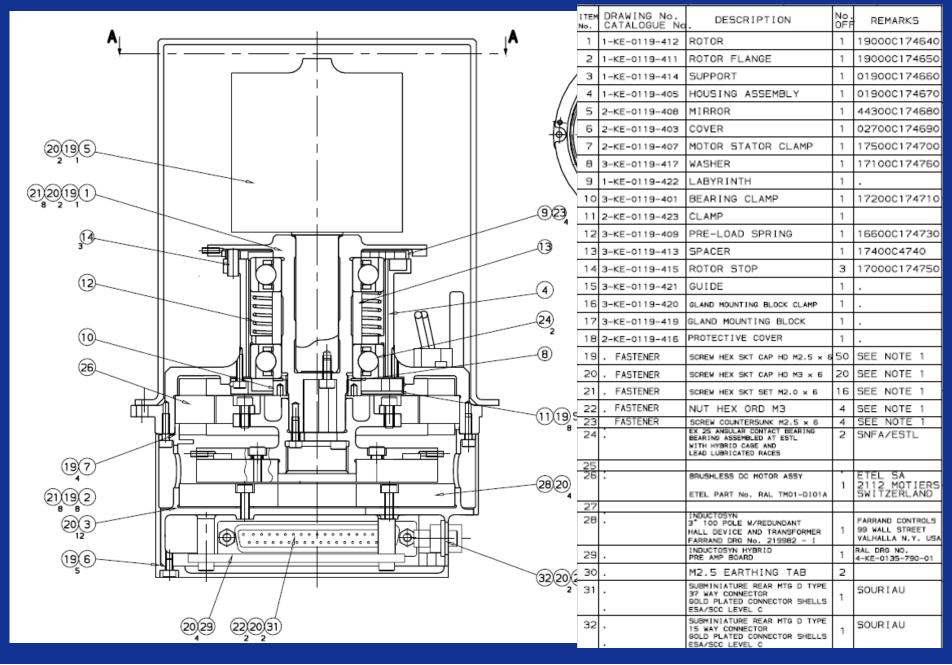
GIST 30, ERB Workshop Ecole Normale Supérieure, Paris 13th – 16th September 2010

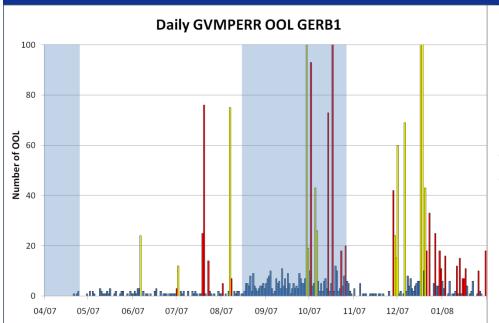
GERB1 Operations Summary

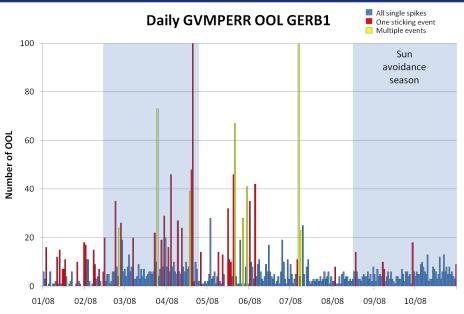


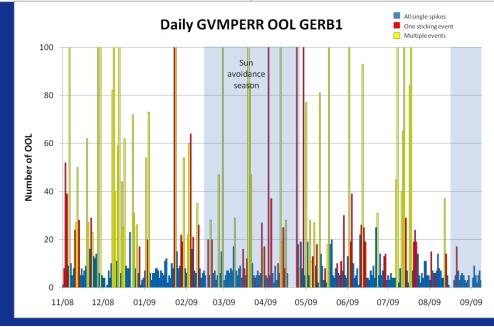
GERB2 Operations Summary

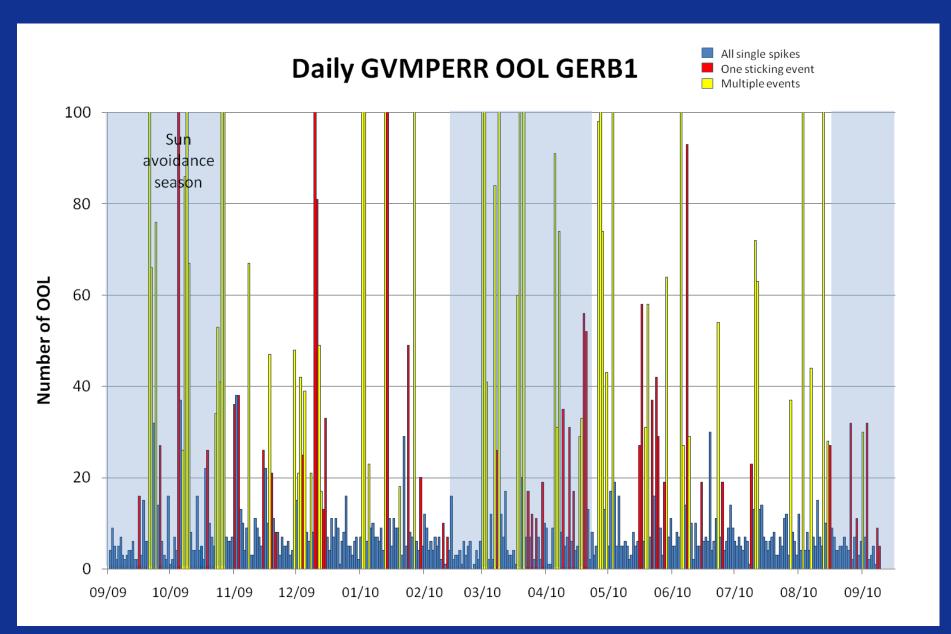




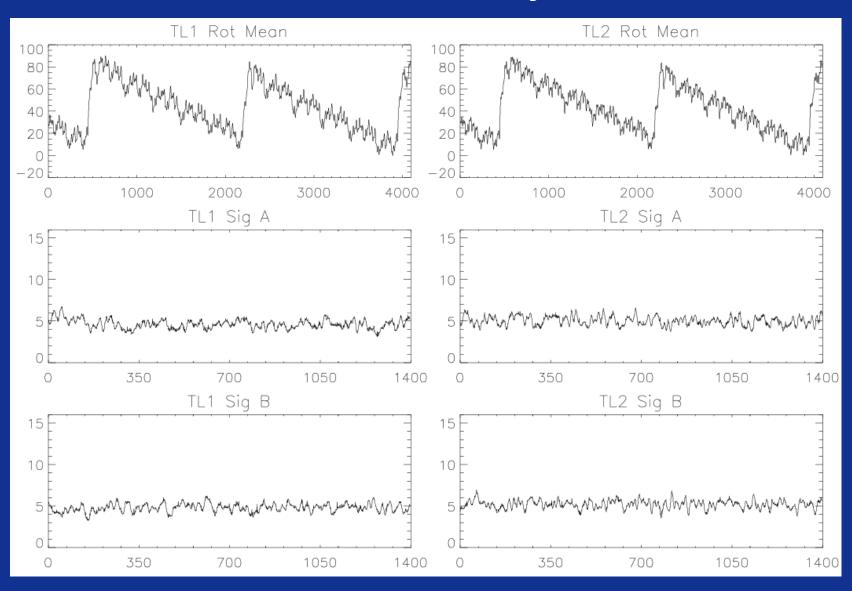




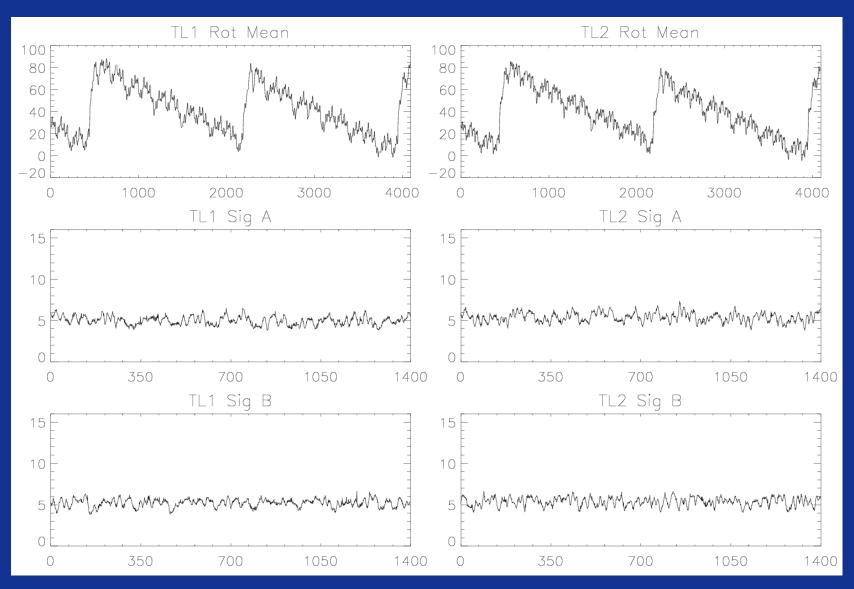




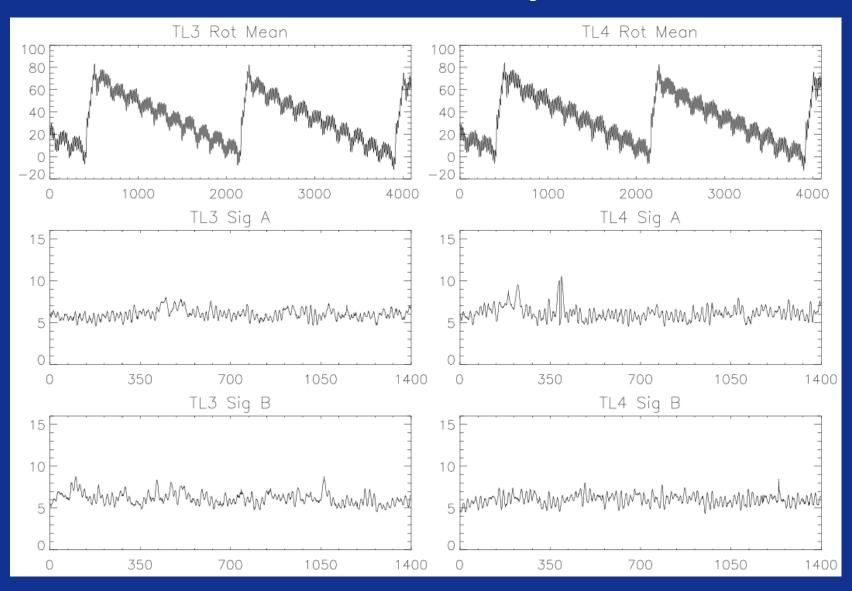
GERB1 TL Test – May 2009



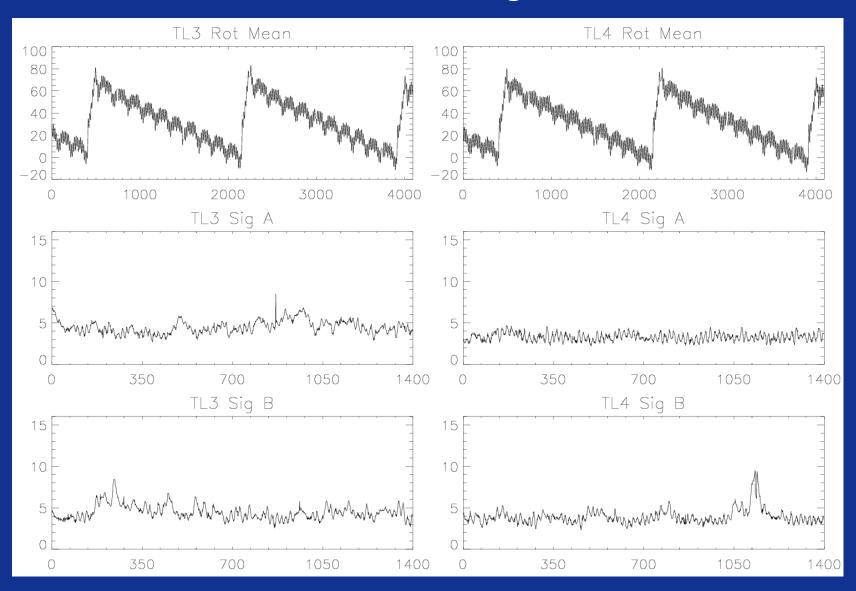
GERB1 TL Test – August 2010



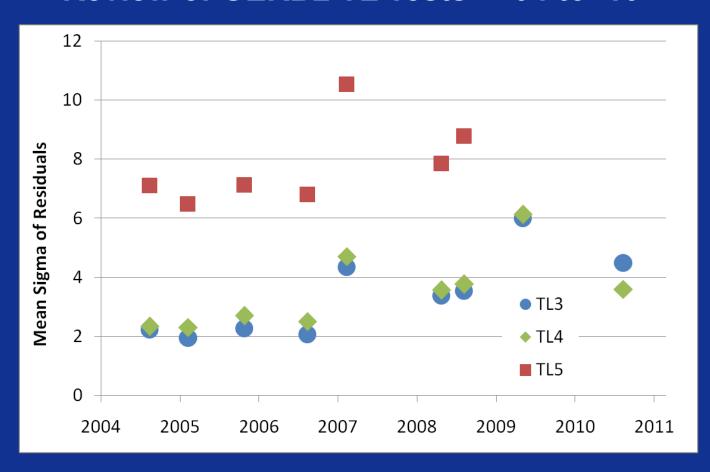
GERB2 TL Test – May 2009



GERB2 TL Test – August 2010



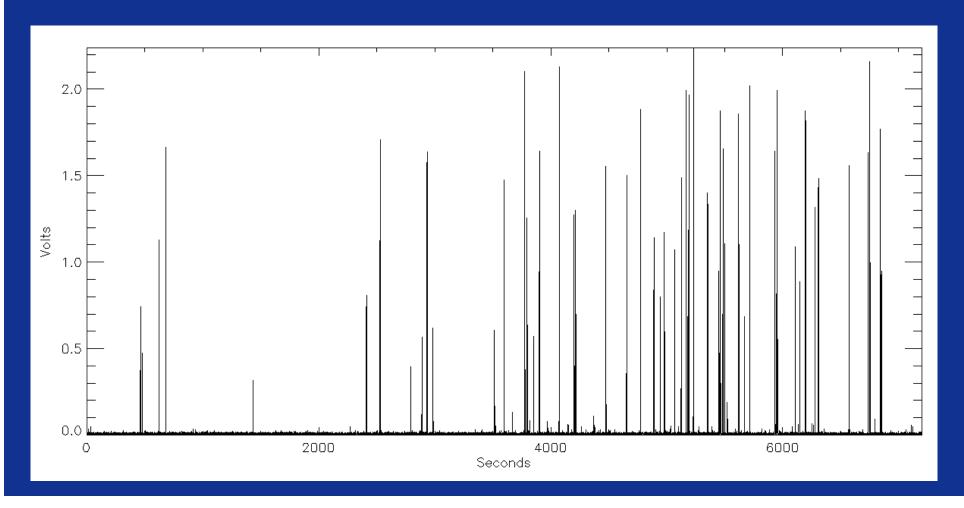
Review of GERB2 TL Tests - '04 to '10



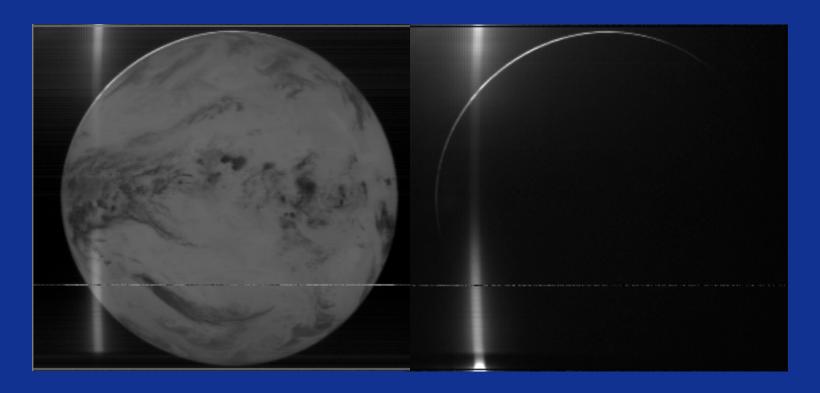
- Changes between 06-07 and 08-09 are in periods of constant mirror rotation. Between 09-10 the mirror was predominantly static.
- Stop testing TL5 due to bearing housing rattle.

TL GERB2 Noise Inversion

Over the course of the two hour torque level test the number of mirror OOL increased significantly. The first hour (TL4) experienced significantly fewer mirror anomalies resulting in lower noise than the second hour (TL3).



GERB Stray Light

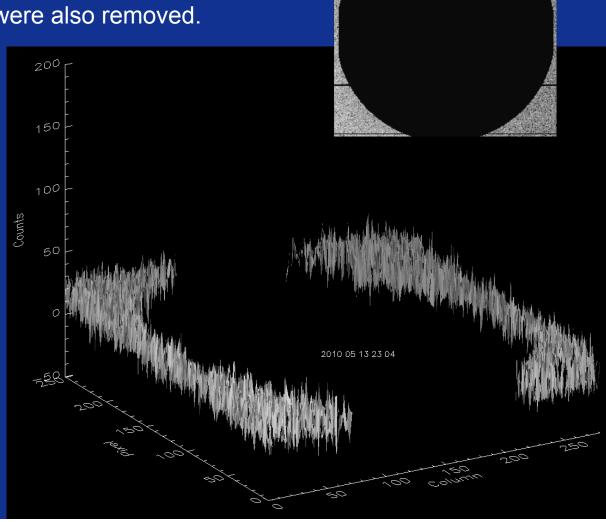


Both GERBs show evidence of stray light. The images shown are from GERB2 just before local midnight. The stray light in this case consists of a bar N-S across the image but more diffuse regions of stray light are present before and after and at approximately midday in the days leading up to and immediately after the Sun avoidance seasons.

Characterising GERB Stray Light

To recover the stray light accurately only the SW, deep space pixels were considered. A circular mask excluded Earth data and noisy pixels were also removed.

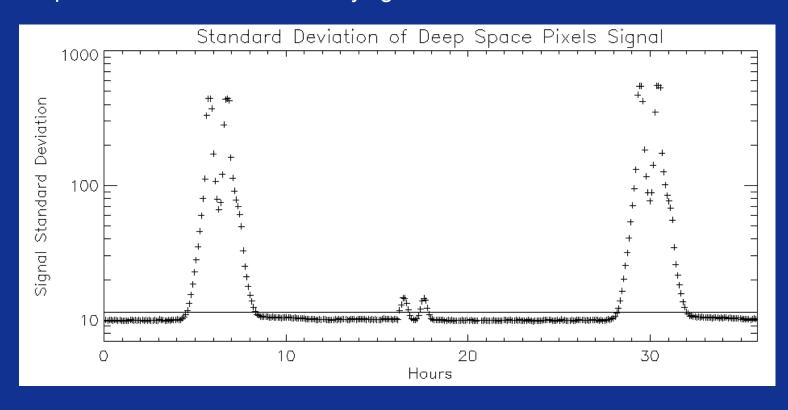
The deep space data plotted shows a typical scene, free of stray light and showing only detector noise.

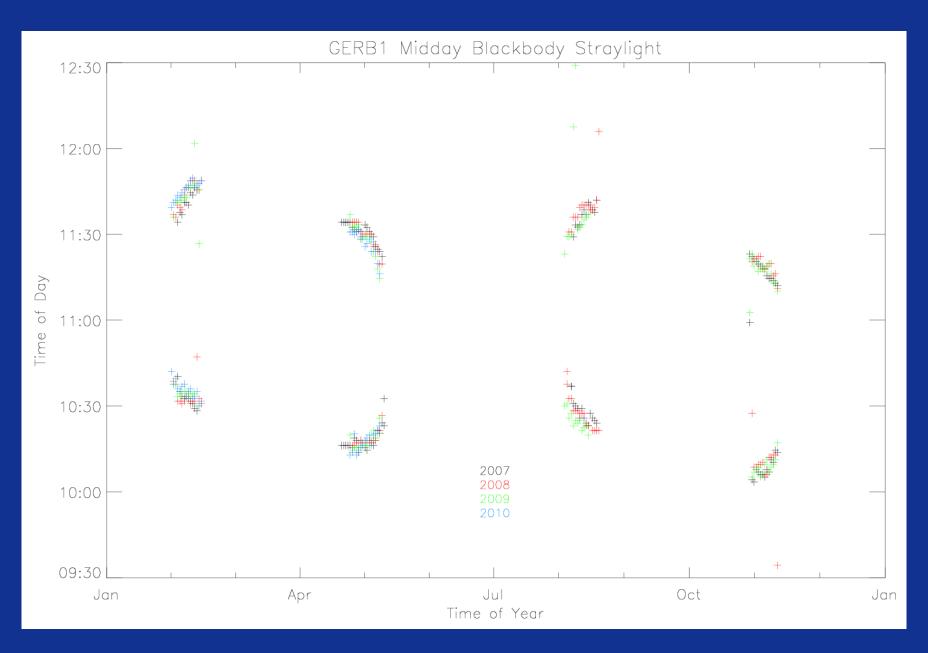


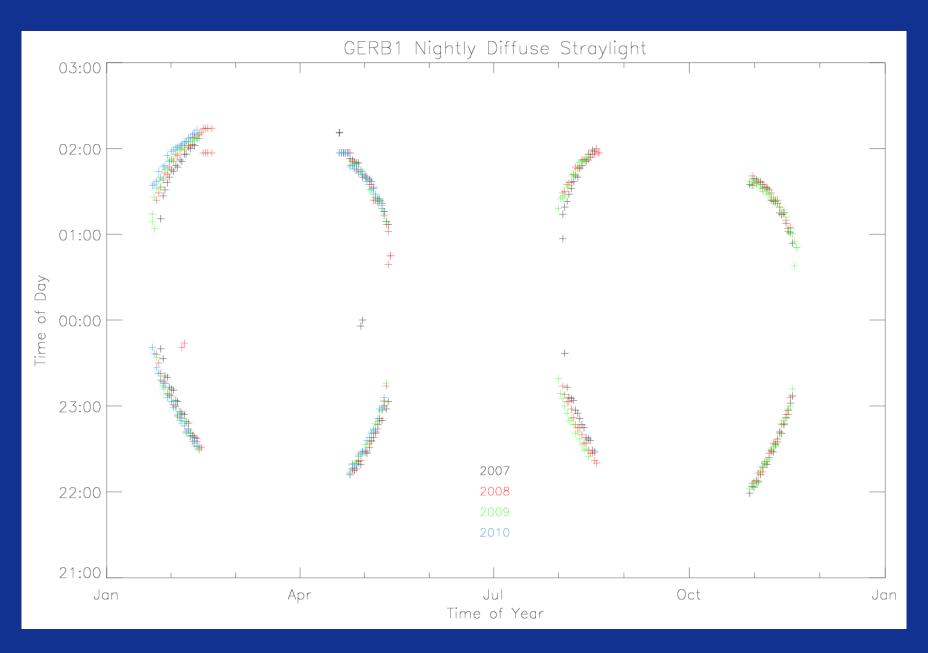
Diffuse Stray Light Threshold

Standard deviations for the deep space pixels of each SW GL0 file over one and a half days centred around midday, 12th February 2010 show clearly the strong signal around midnight and the smaller problem at midday.

A threshold of 115% of the background level was found to give a indicator for the presence of all forms of stray light.



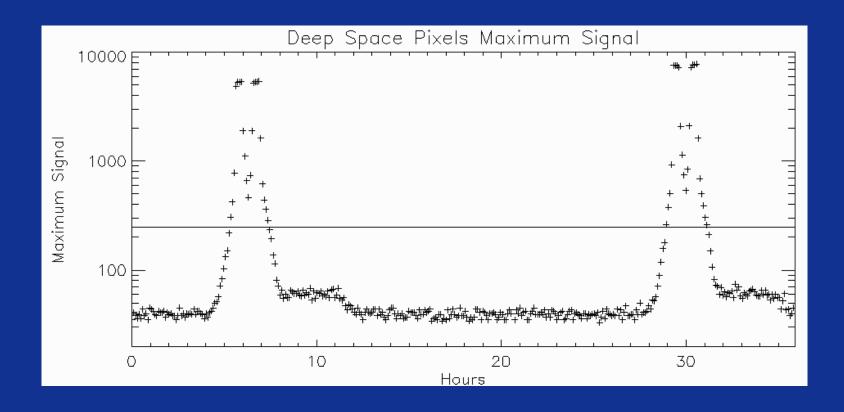


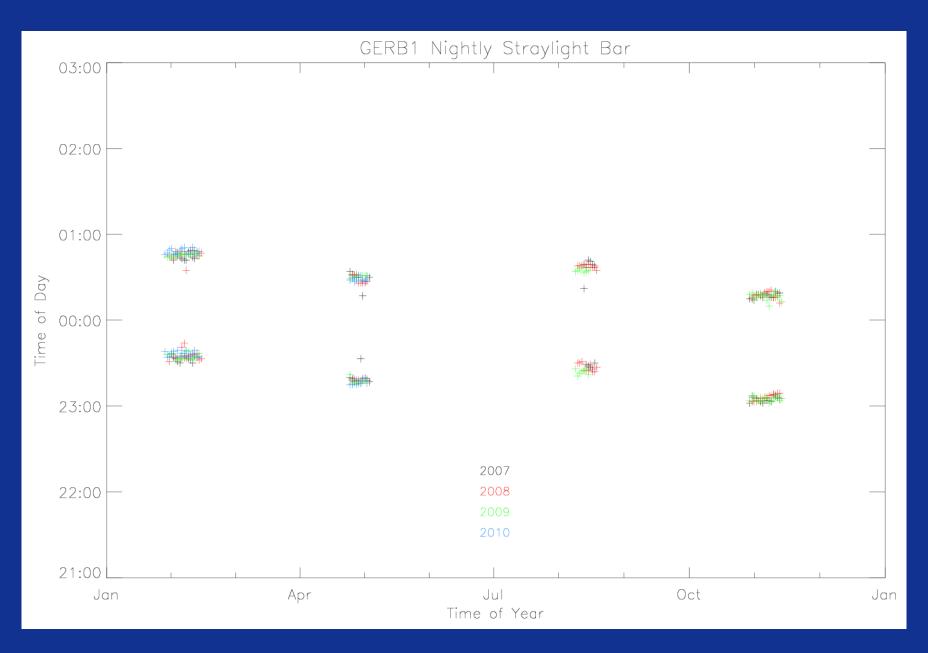


Bar Stray Light Threshold

The nightly bar passing W-E across the GERB image was is shown from data taken centred around midday on the 12th Feb 2010.

A signal threshold of 250 was identified by inspection to separate the appearance of the bar structure from high levels of diffuse stray light.





Summary

Mirror bearing performance

GERB1

• The number and duration of mirror OOL remains stable at a level which renders extended data taking in the SA season too dangerous.

GERB2

- Extended running of the ageing mirror bearing shows increasing noise although this seems to improve after long periods of inactivity.
- There is a small amount of mirror data so recent trends are tentative.
- Continue running GERB2 sparingly, test TLs again post SA season.

Stray Light

- Clear trends identified in the appearance of the GERB1 stray light.
- Extend analysis to GERB2 edition region (Feb 2004 May 2007.)